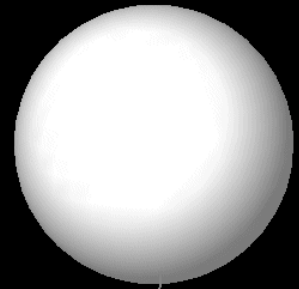


GUIDED MARS BALLOON PLATFORMS



**Presentation to International Planetary
Probe Workshop, June 28, 2006**

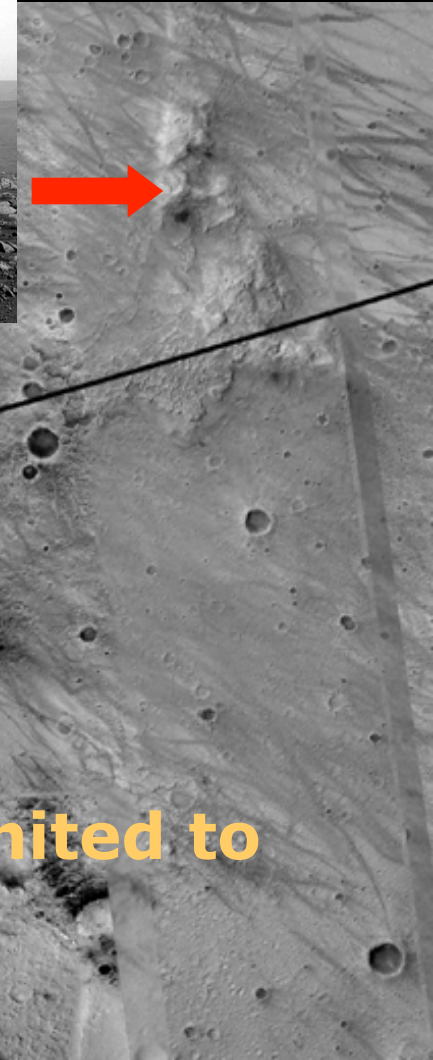
**A. A. Pankine, K. T. Nock, N. C. Barnes
Global Aerospace Corporation**



MARS ROVERS ARE A GREAT SUCCESS...



Spirit view from Husband Hill
summit (NASA/JPL)



... but their range is limited to
several miles

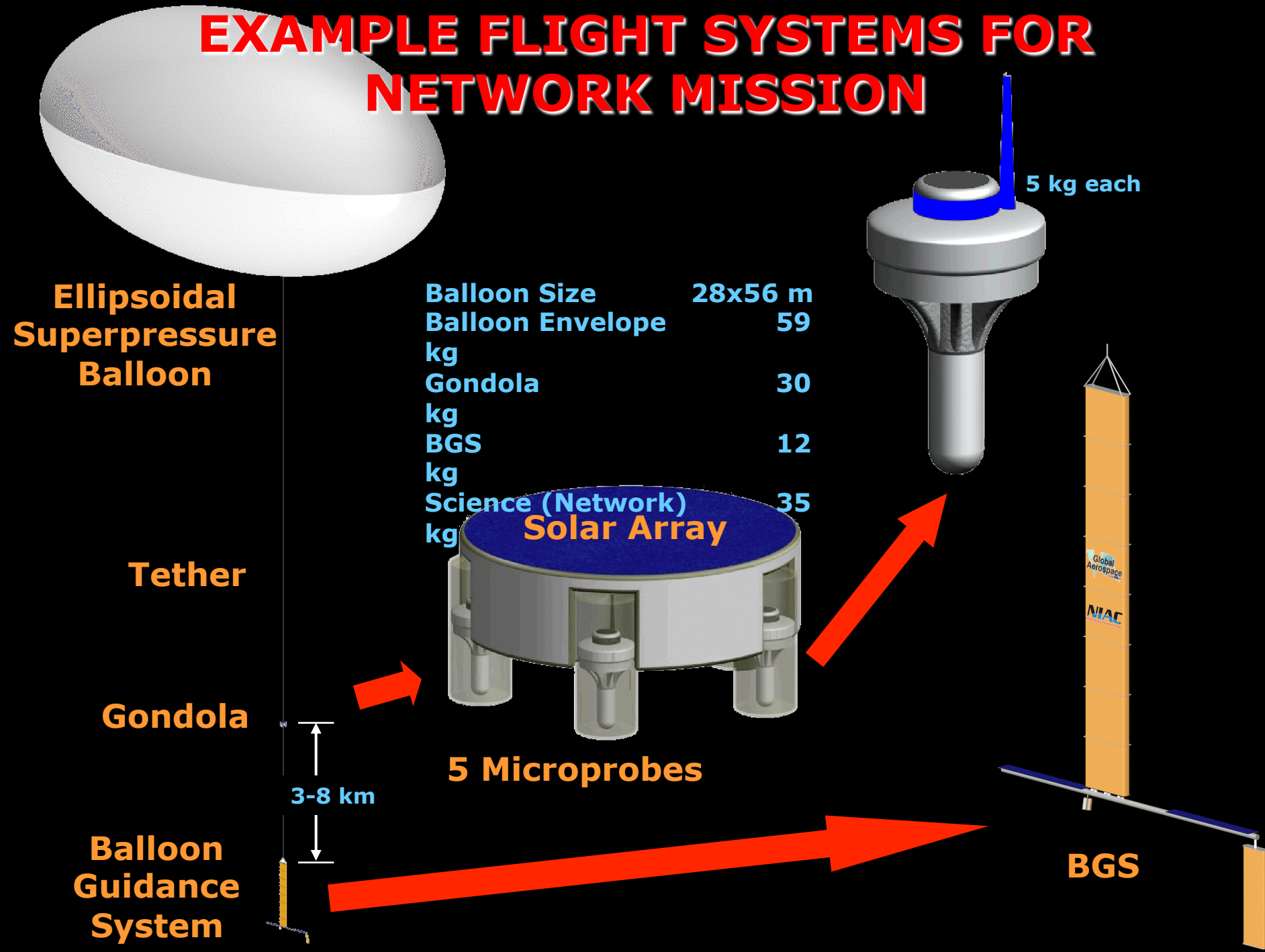
Columbia Hills surroundings (NASA/
JPL/MSSS)

A NEW PLATFORM FOR EXPLORATION

- Orbiters are not *in situ*
- Landers do not move
- Rovers have very limited range
- Airplanes or gliders last for just a few hours
- Airship propulsion makes them heavy and difficult to deploy
- Free balloons are totally at the mercy of the winds

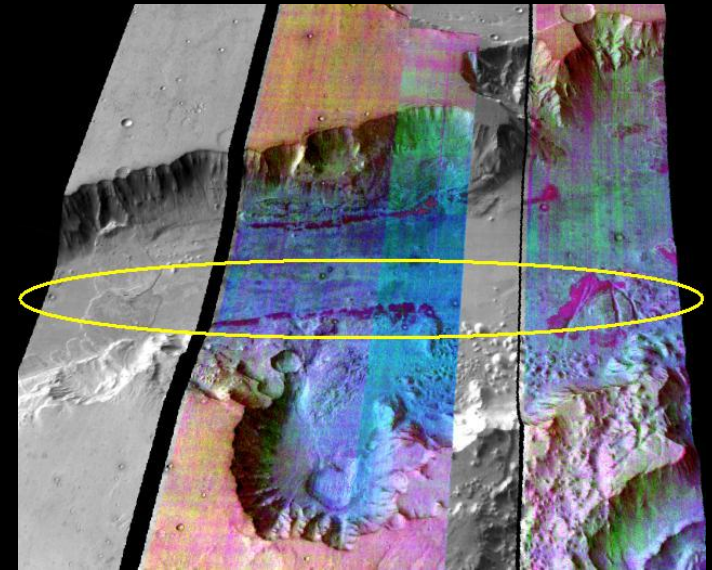
*Guided, long-duration
balloon platforms
have global reach*

EXAMPLE FLIGHT SYSTEMS FOR NETWORK MISSION

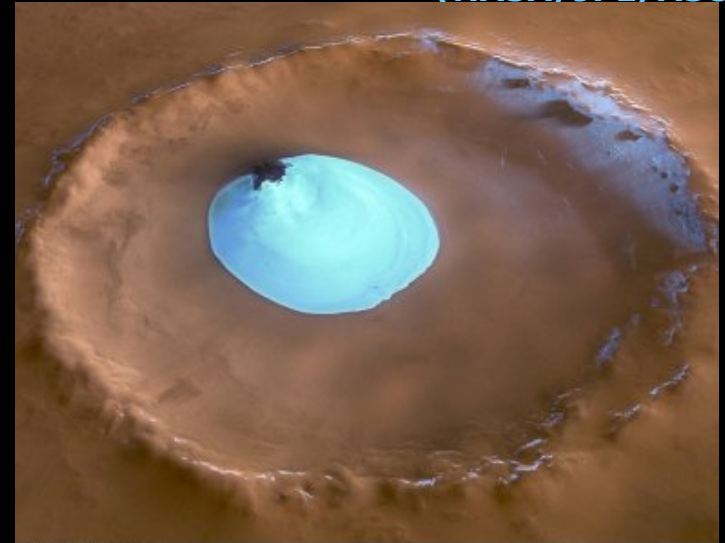


EXPLORATION CAPABILITIES

- Global planetary coverage
- Heavy, power-intensive payloads (90 kg and 200 W in 3 to 10 years, 170 kg and 400 W >10 years)
- Long flight duration: 700 days (1 Mars year)
- Autonomous navigation and guidance, target acquisition
- Targeted overflight of surface sites and precise delivery of science probes
- High-resolution imaging (1-10 cm visible & 0.1-10 m IR), elemental, magnetic and gravity observations and surveys
- In situ atmospheric chemistry and circulation
- Landing sites reconnaissance, navigation beacon emplacement



Olivine outcrop and DS-2 landing ellipse
(NASA/JPL/ASU)

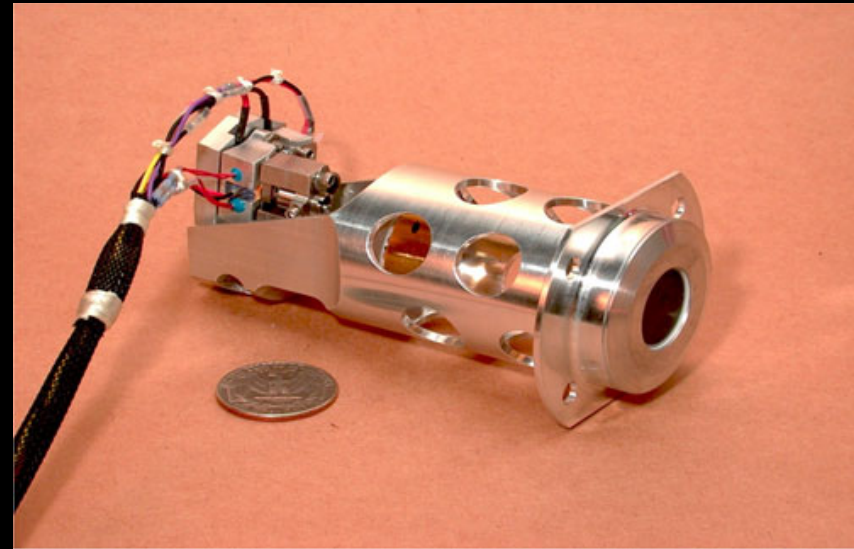


© ESA/DRFU Berlin (D. Neukum)

Water ice lake inside a crater on Mars
(ESA)

SEARCH FOR ORIGIN OF ATMOSPHERIC METHANE

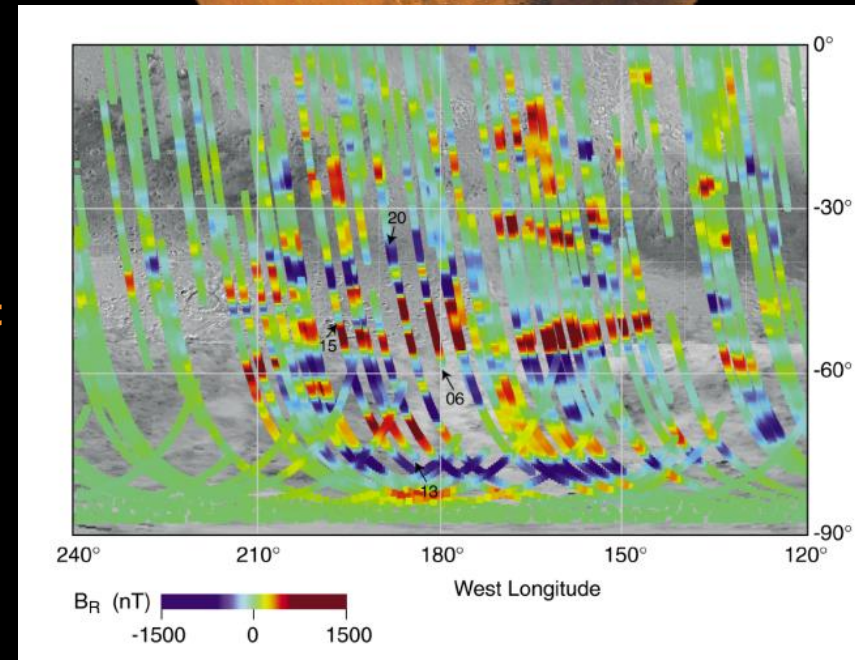
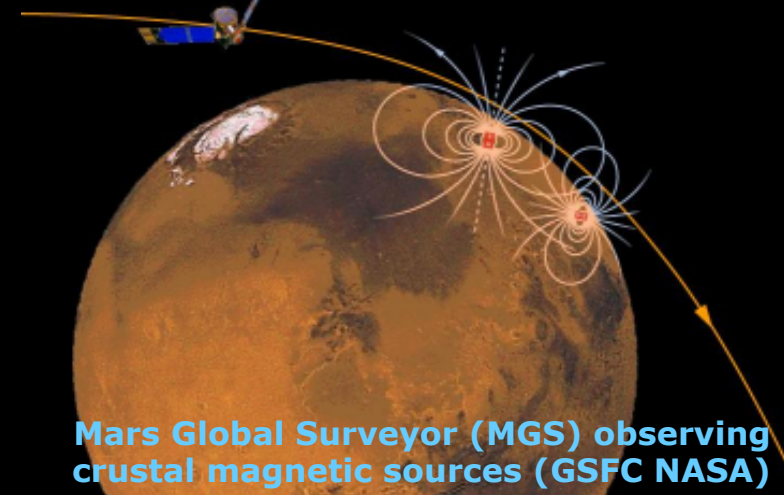
- Methane-making organisms discriminate between isotopes as they feed on a global reservoir of CO₂
- Measure the C¹²/C¹³ ratio in the methane.
- If it is different from the isotope ratio in the CO₂, it would offer strong evidence for a biological source.
- Guided balloons enable planetary-wide search for the sources of methane



Tunable Laser Spectrometer for Atmospheric and Sub-surface gas measurements on Mars (NASA JPL)

CHARACTERIZE CRUSTAL MAGNETIC ANOMALIES

- MGS discovered strong crustal magnetic field anomalies
- However, orbital measurements
 - Lack required resolution to resolve origin of anomalies, and
 - The solar wind obscures weak fields
- Understanding these anomalies will
 - Provide clues on crust evolution
 - Advance understanding of Martian extinct dynamo
 - Possibly identify subsurface structures that harbor life
- Balloon platform with an array of magnetometers can
 - Enable high-resolution observation of crustal magnetic anomalies
 - Enable detection of weak anomalies via gradient measurements



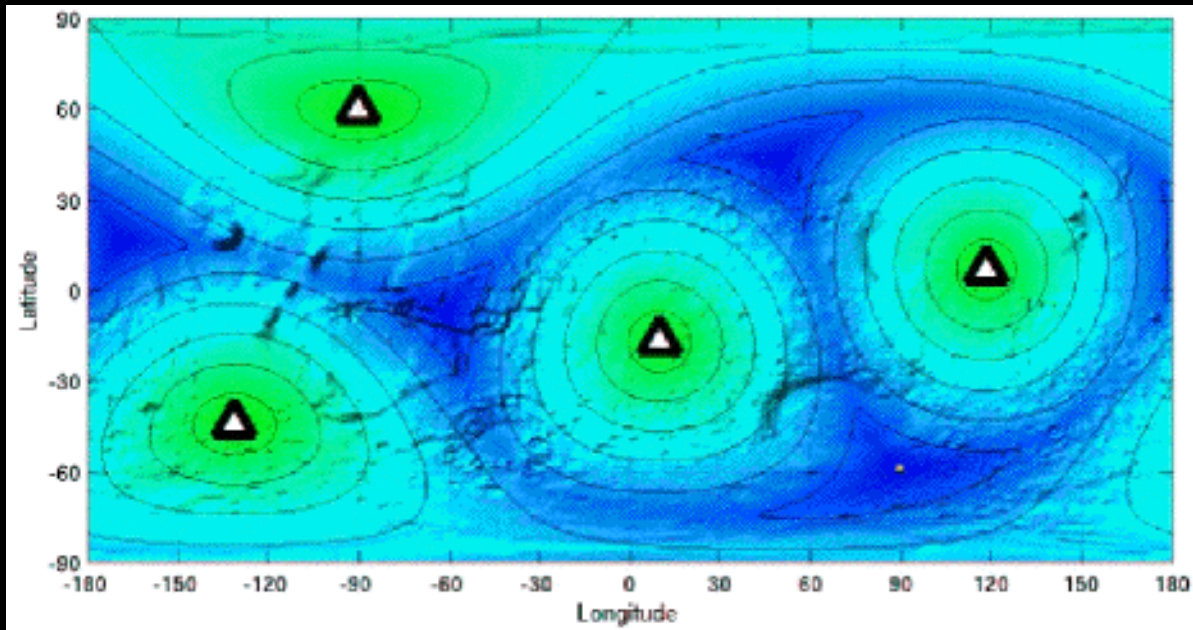
Map of crustal magnetic anomalies on Mars (GSFC NASA)

EMPLACEMENT OF SURFACE NETWORKS ON MARS

A single guided balloon platform can

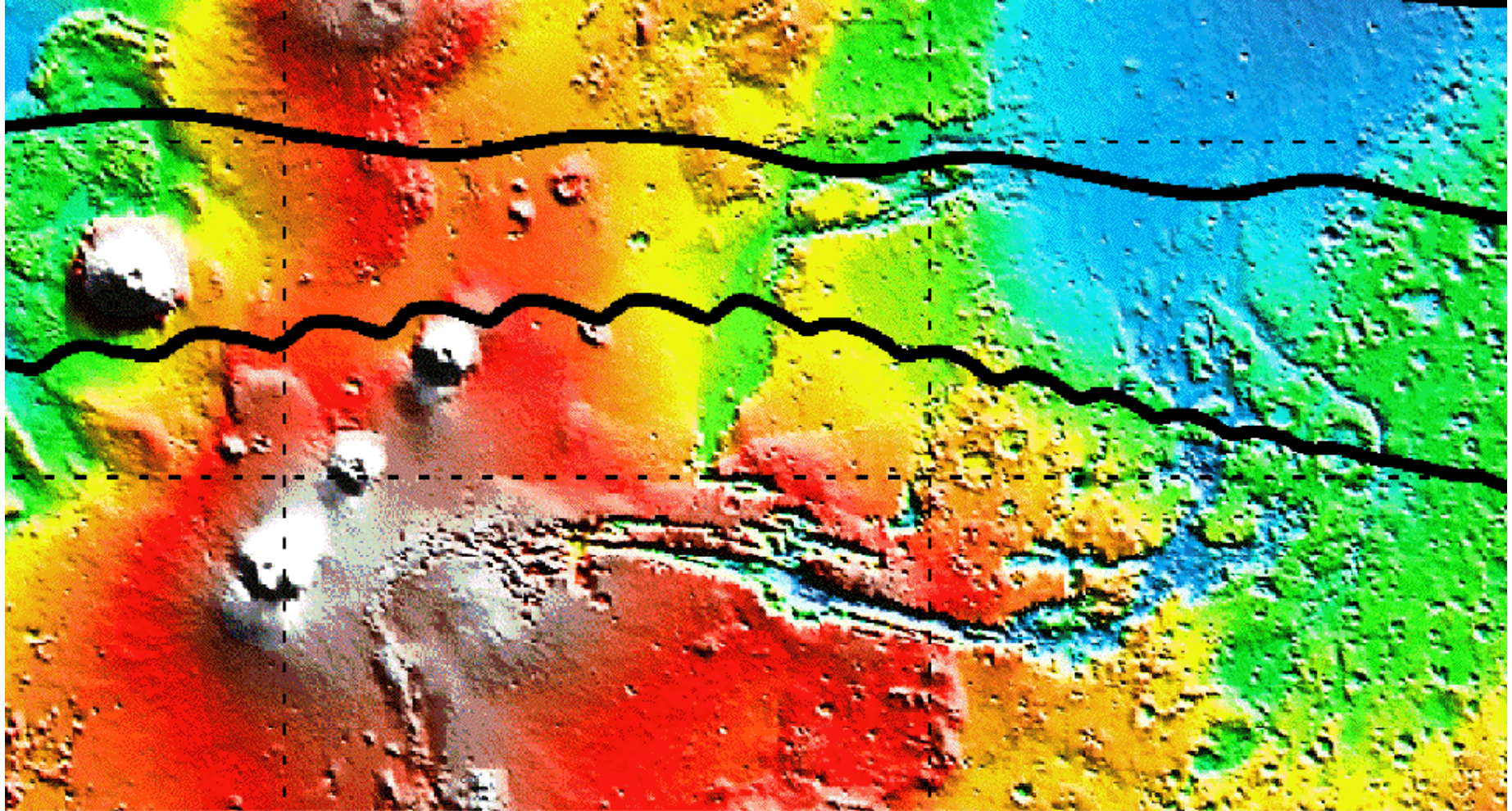
- Carry tens of mini-labs
- Deploy meteorological & seismological networks

Tetrahedron Seismometer Network

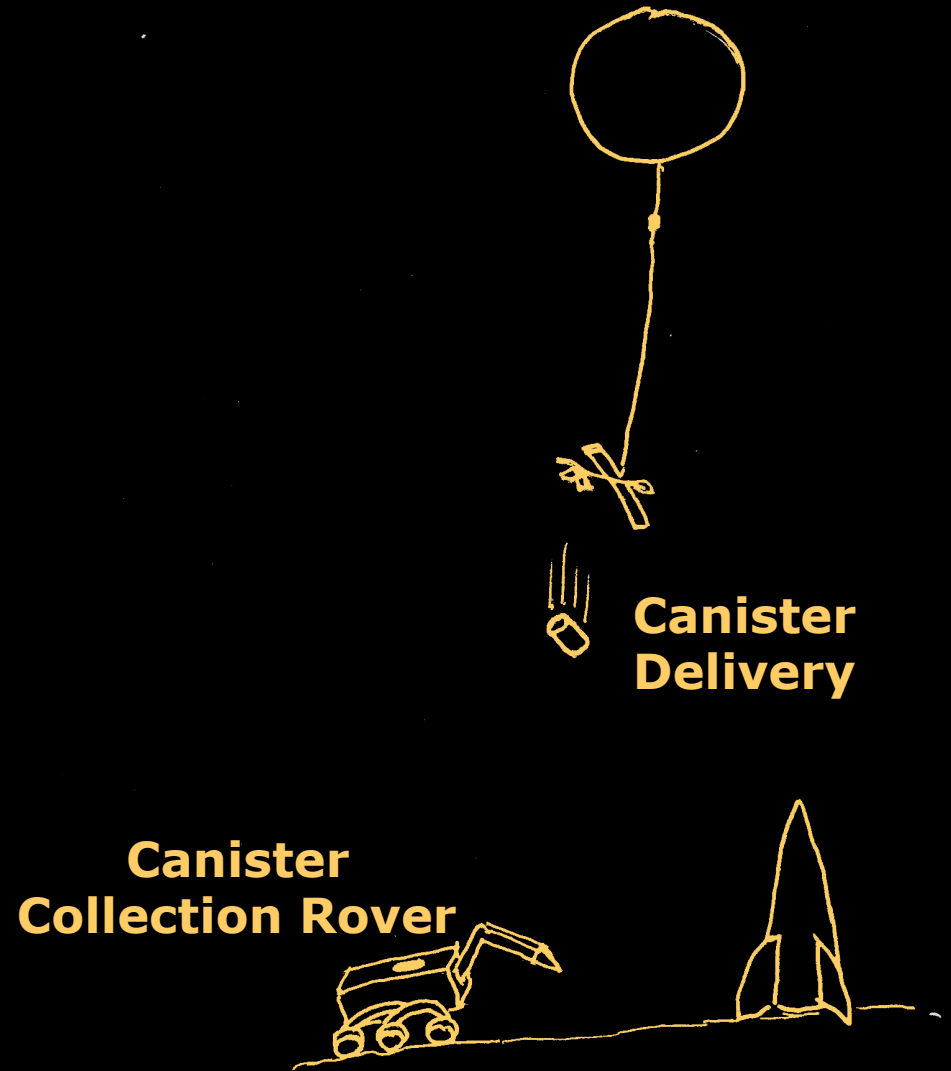
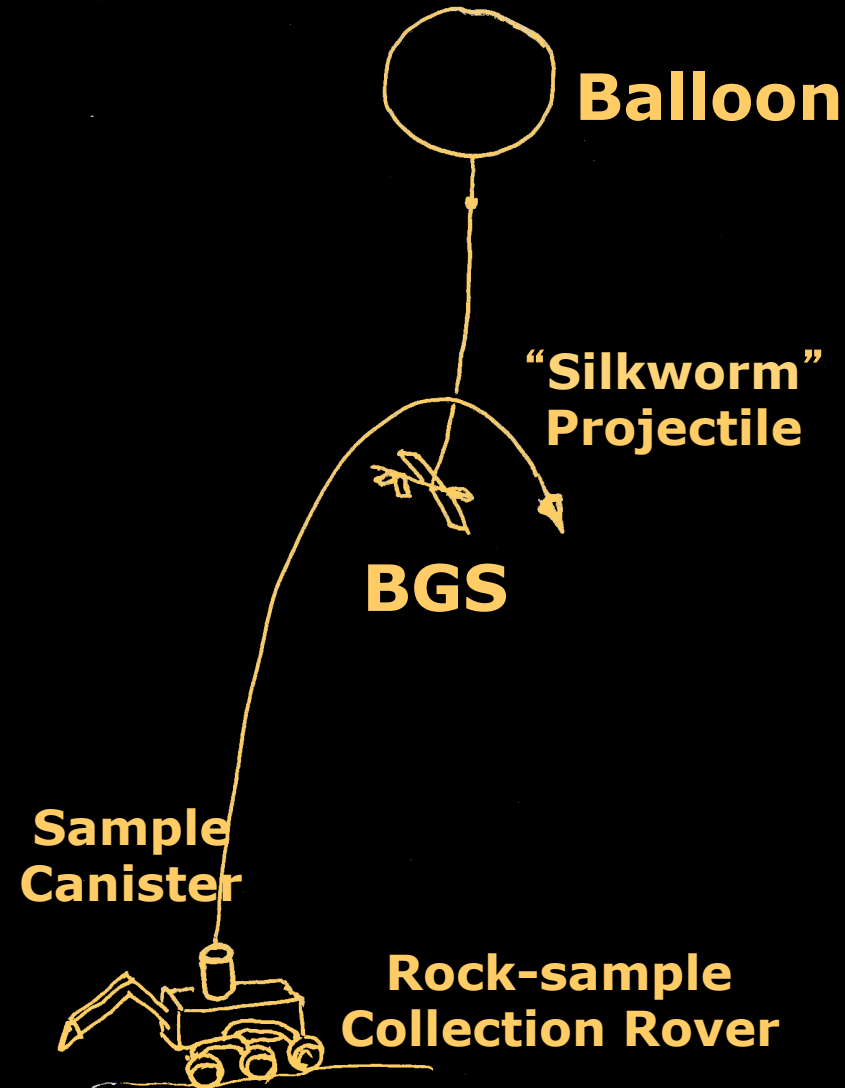


**Mars Meteorological
and Seismometer
Probe**

SAILING ACROSS THE MARTIAN EQUATOR



SAMPLE RETURN ASSIST



THE OTHER SIDE OF MARS

SAMPLE RETURN ASCENT VEHICLE

SUMMARY

Guided Mars balloon platforms

- **Extend the reach of and work synergistically with orbiters, rovers, and landers**
- **Can be a relatively low-cost component of future Mars exploration**
- **Enable revolutionary new planetary exploration capabilities at Mars**

MARS WIND PROFILE AND BGS

- **Winds vary with altitude**

- Balloon at 10 km

Density = 0.00645 kg/m^3

- Wing at 3 km

Density = 0.01186 kg/m^3

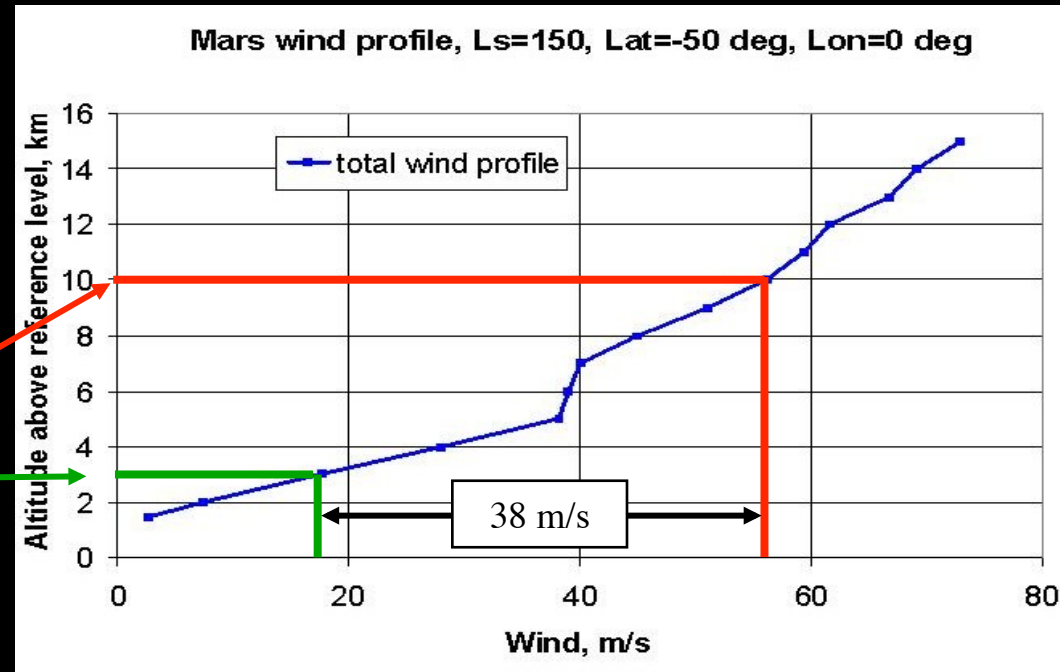
- Relative wind velocity
~ 38 m/s

- **Wing generates lift force**

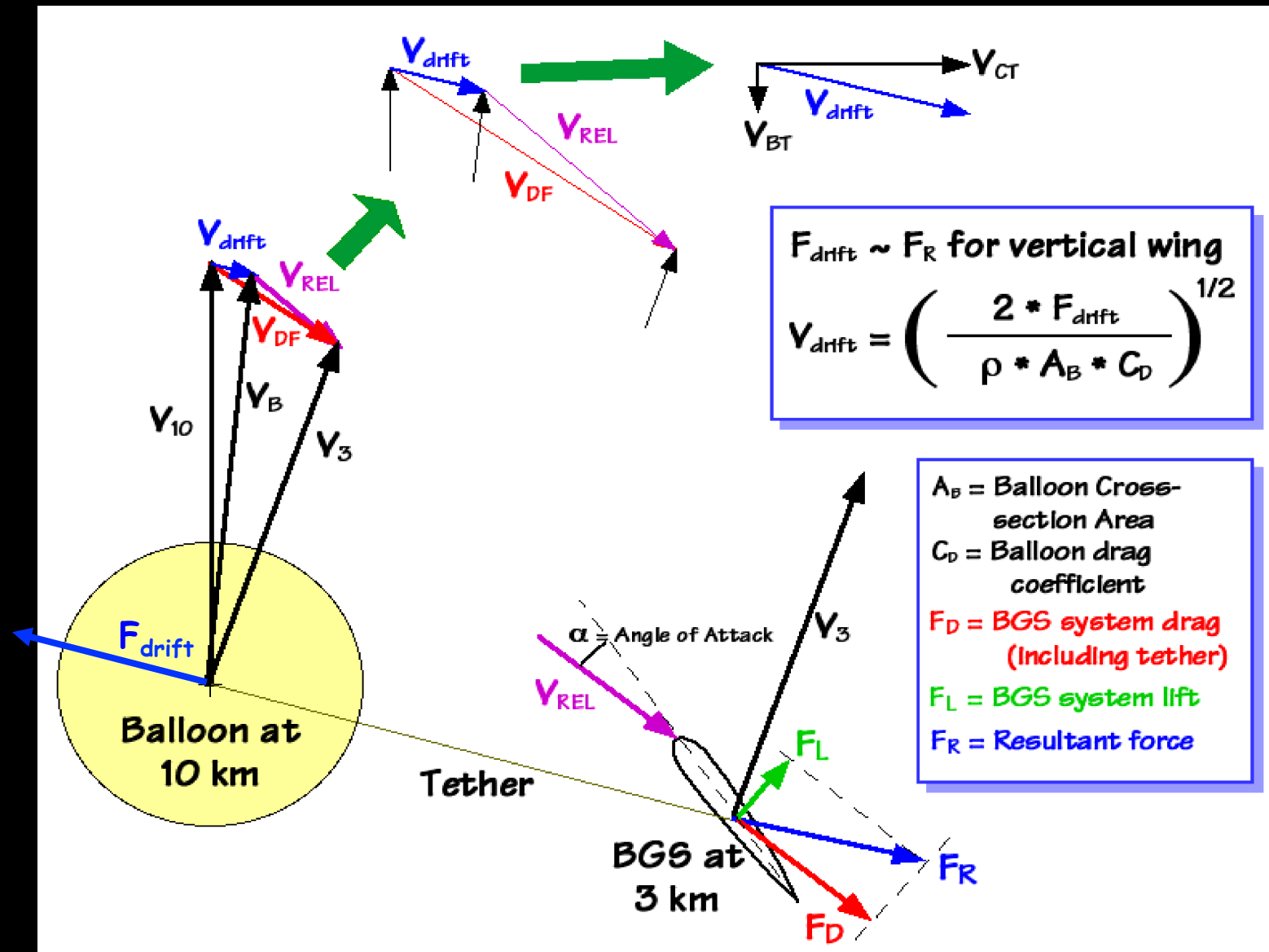
- “Lift” force is horizontal
- force is transmitted by tether to balloon
- balloon drifts relative to local air mass
- balloon drag \approx wing lift

- **Wing is in denser air than balloon**

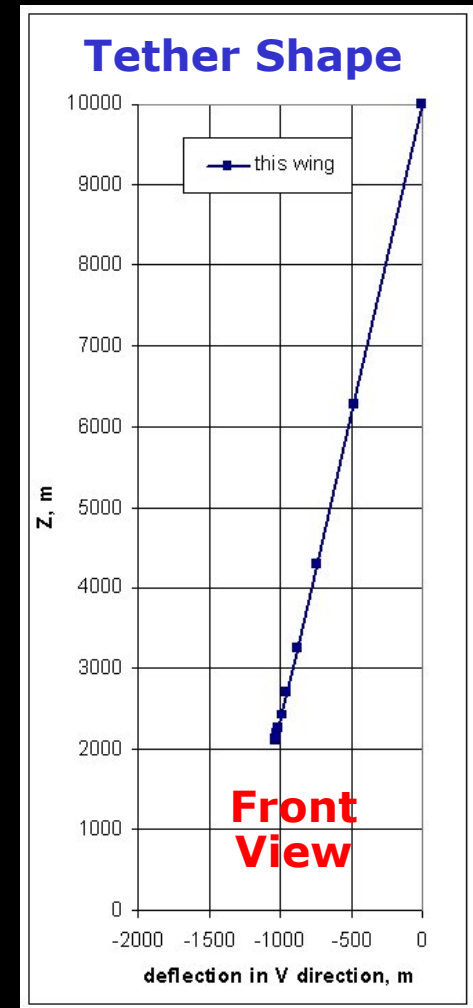
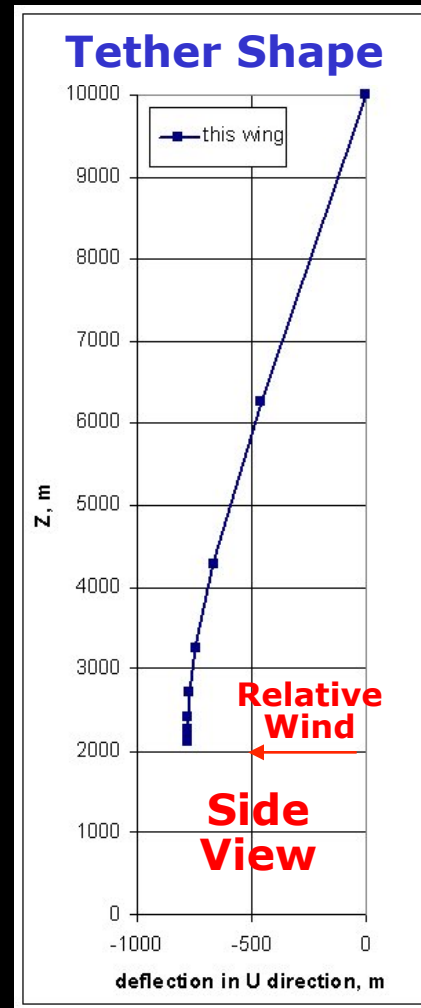
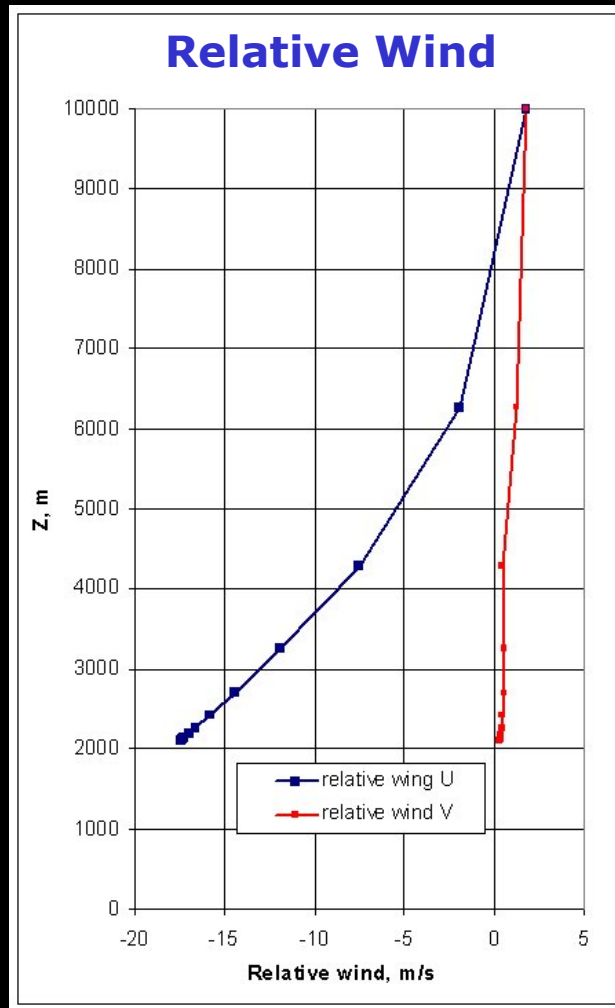
- 3 km : 10 km (1.84x)
- equivalent wing area increased relative to balloon



BGS OPERATION

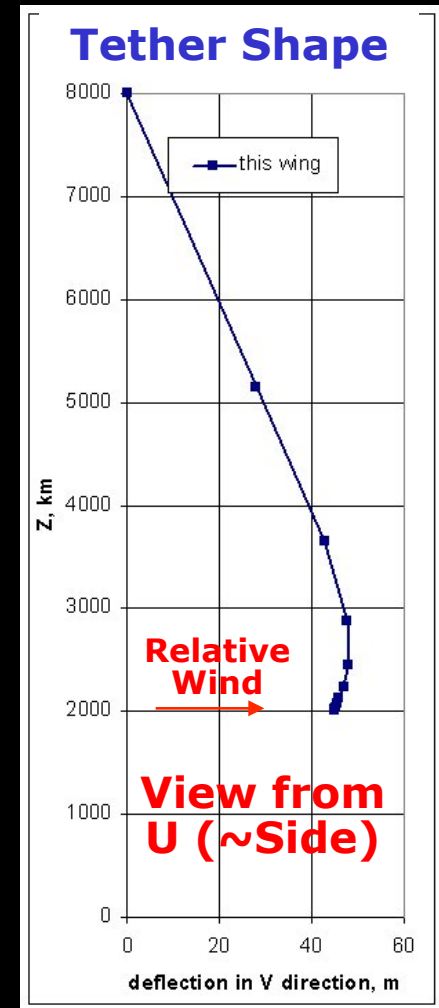
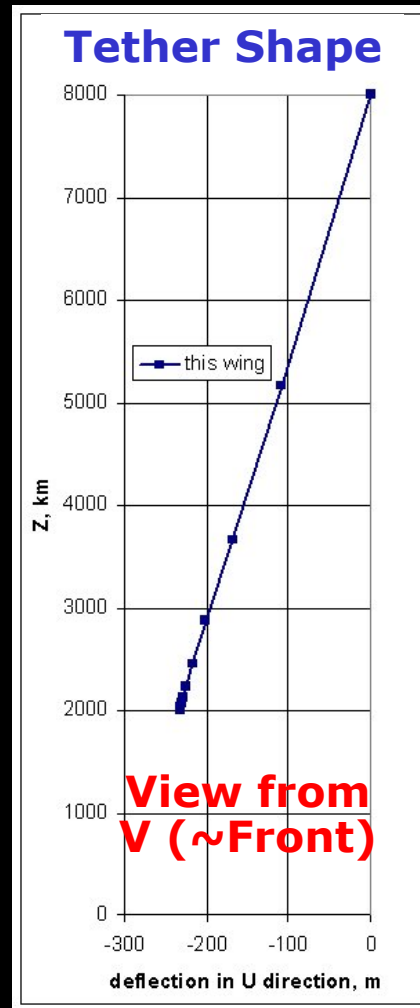
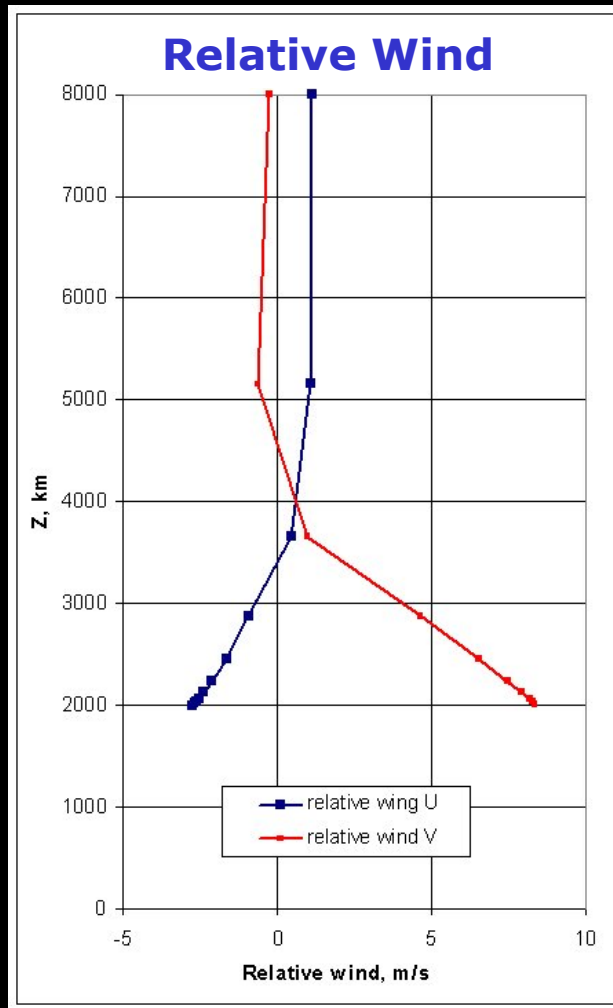


BGS PERFORMANCE ANALYSIS: STRONG ZONAL FLOW



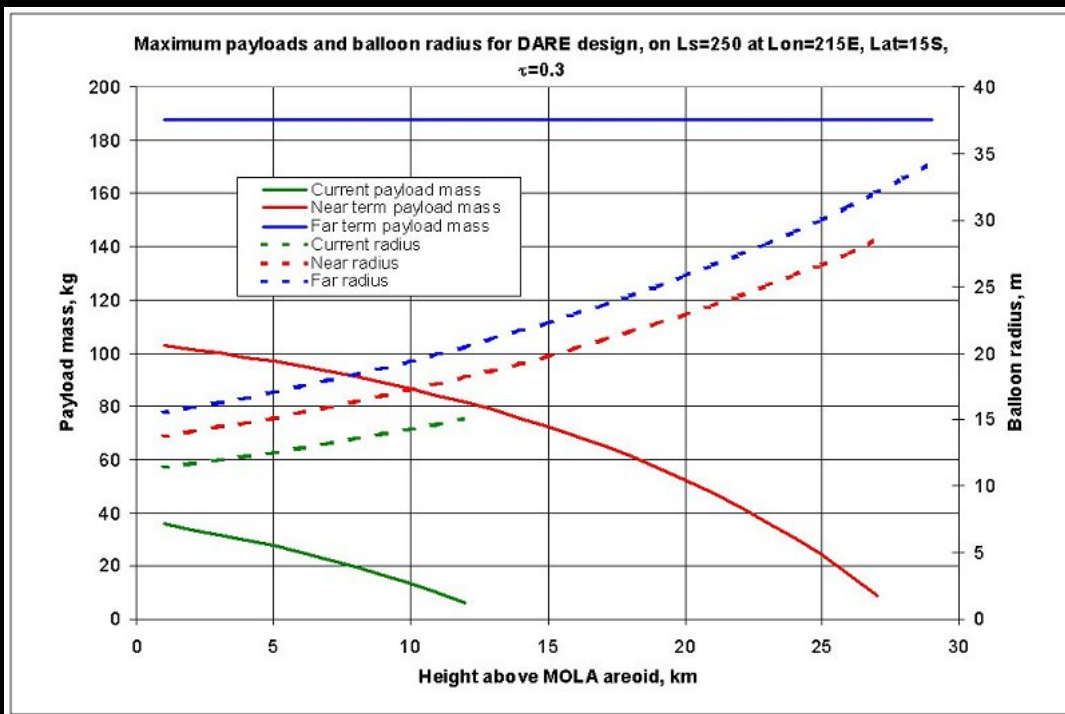
- BGS $V_{rel} = -18$ m/s
- BGS $\Delta V = -1.8$ m/s

BGS PERFORMANCE ANALYSIS: WEAK FLOW

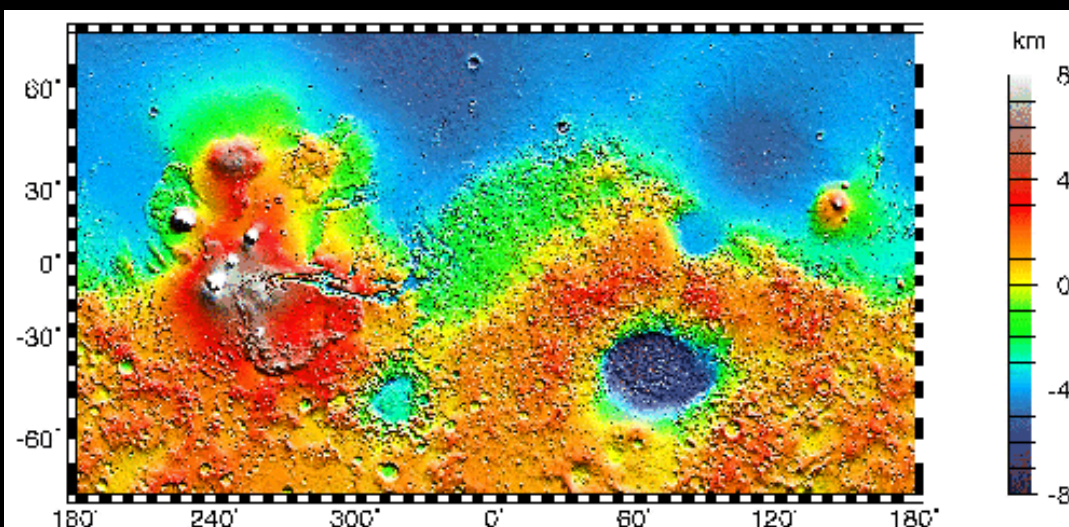


- BGS $V_{rel}=9$ m/s
- BGS $\Delta V =1$ m/s

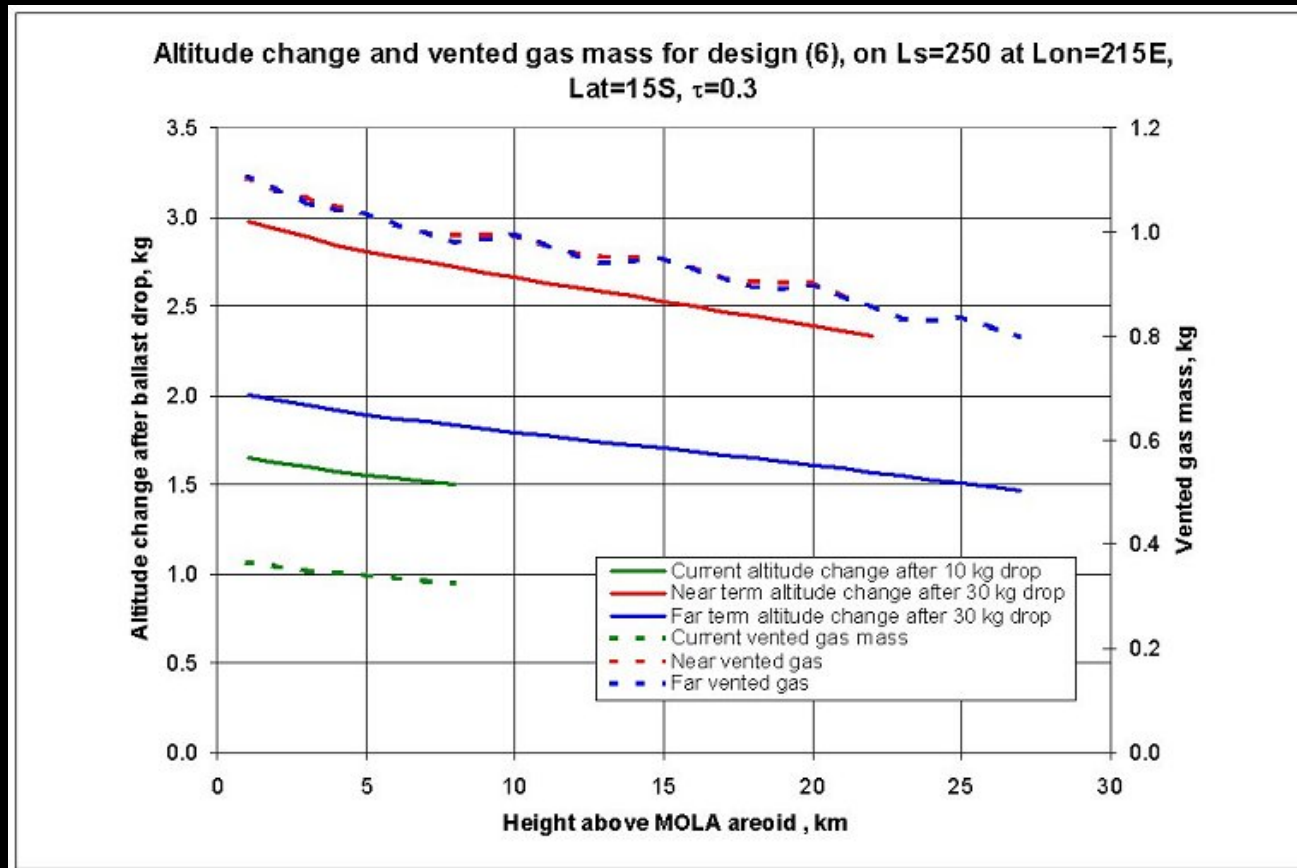
PAYLOAD VS. ALTITUDE



- Height of atmospheric density levels lower by 4 km in dusty atmosphere
- Balloon to float 2-3 km above southern highlands in dust storm
- 6 km at $\tau = 3$
- $M=87$ kg, $R=17.2$ m
- Altitude of 10 km at normal conditions



ALTITUDE CHANGE AFTER PROBE RELEASE



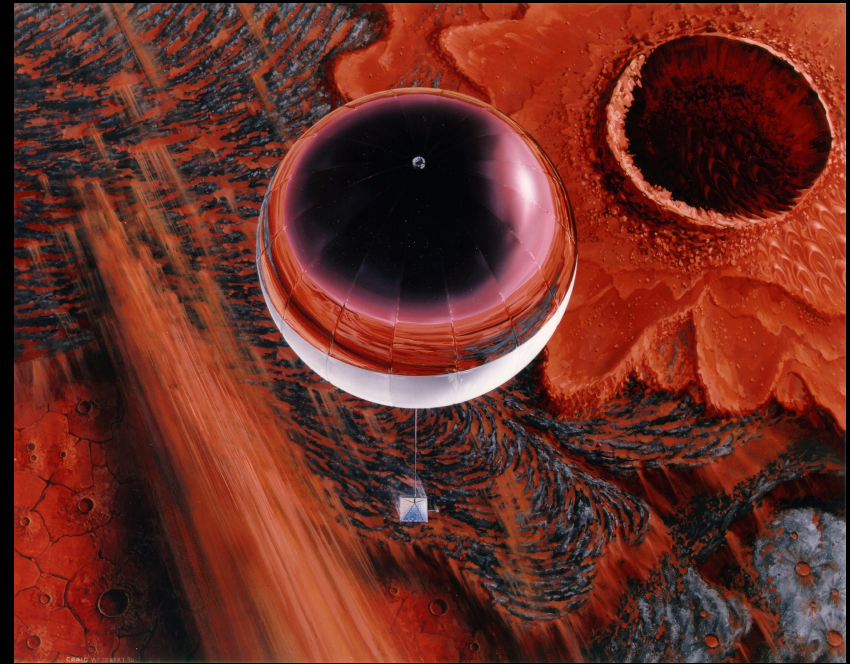
- Releasing 30 kg of probes raises altitude by 3 km
- Increase in super-pressure can be relieved by venting 1 kg of gas (out of 8 kg)

KEY TECHNOLOGIES

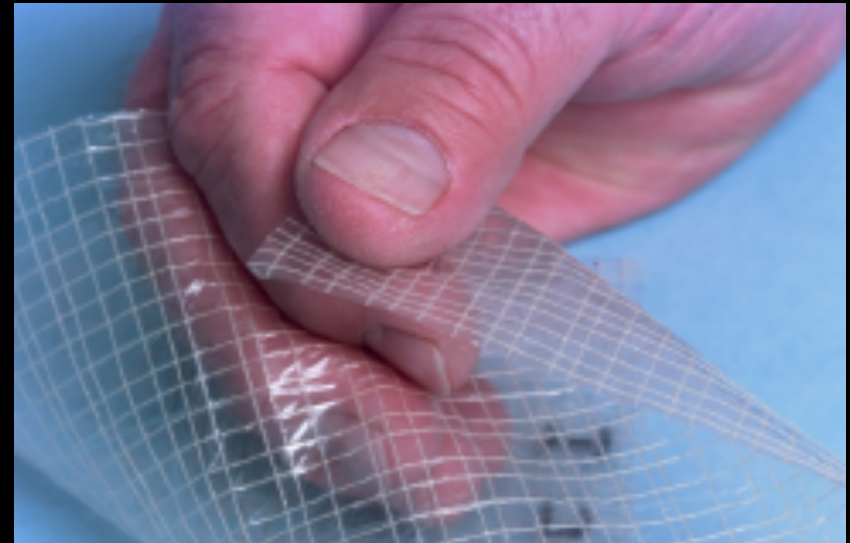
- **Advanced Balloon Materials**
- **Balloon Guidance System (BGS)**
- **Entry, Descent and Inflation (EDI)**
- **Autonomous Navigation & Guidance in Mars winds**
- **Mars Balloon performance modeling**

MARS BALLOON

- **Low-mass high-strength envelope material**
 - composite material
 - 1- μm Mylar/38-Denier PBO thread/3- μm PE film
 - areal density of 0.012 kg/m²
 - Nano-tubes fabric in future?
- **Superpressure sphere**
- **Aluminized top, white bottom to prevent CO₂ condensation**



Mars balloon concept

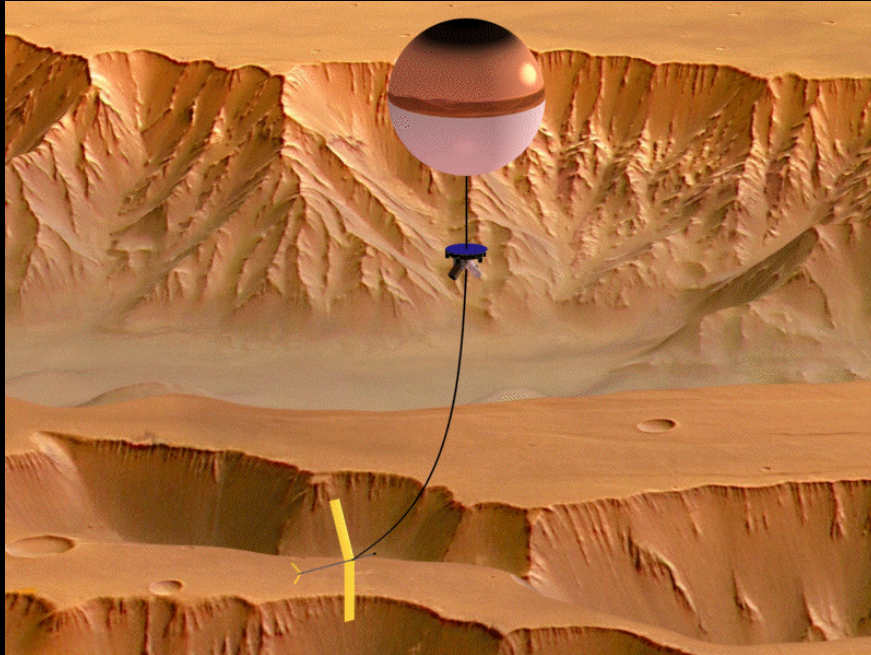


Composite Mars balloon material¹⁹

EDI

- **Mach 2 parachute at 13-28 km**
- **Careful balloon envelope deployment**
- **Cryogenic H₂ inflation - near-term technology**
 - **Inflation gas (8.3 kg H₂)**
 - **Combustion gases (0.14 kg H₂ and 1.14 kg O₂)**
 - **Heat exchanger**
- **Float at 8 km after 20 min**
- **Make-up gas for >100 day missions**
 - **~1.6 kg H₂ per year minimum**
 - **Collect and crack ambient H₂ O for H₂ (10⁻⁴ kg H₂ O per kg of atmosphere)**

AUTONOMOUS NAVIGATION AND GUIDANCE



- **Navigation**
 - star camera
 - trajectory forecast
 - orbiter communication
 - surface feature recognition
 - control algorithms
- **Target acquisition**
 - onboard target database
 - surface feature recognition
 - targets of opportunity
 - camera attitude sensor and controller